

Perturbation methods



2018/6/7

Outline

- Basic principle
- Degenerate case
- Stark Effect
- Fine structure
- Hyperfine structure

Time independent

- Assume that the Hamiltonian is a sum of two terms,

$$H = H_0 + \lambda H_1$$

- Let $\{|\Phi_n\rangle\}$ be a complete set of eigenstates of the unperturbed Hamiltonian H_0 with energy eigenvalues $E_n^{(0)}$

$$H_0 |\phi_n\rangle = E_n^{(0)} |\phi_n\rangle$$

- The eigenstates $\{|\psi\rangle\}$ and eigenvalues $\{E_n\}$ of the complete Hamiltonian H

$$H |\psi_n\rangle = E_n |\psi_n\rangle$$

To express H with a matrix

- choose basis of $|\Phi_n\rangle$
- The diagonal elements are original eigenvalues

$$H_{mn} = \langle \phi_m | H | \phi_n \rangle = E_m^{(0)} \delta_{mn} + \langle \phi_m | \lambda H_1 | \phi_n \rangle$$

Non-degenerate case

- unperturbed results

$$\lambda \ll 1 \quad |\psi_n\rangle \simeq |\phi_n\rangle \quad E_n \simeq E_n^{(0)}$$

- To express $|\psi\rangle$ using complete set $|\Phi_n\rangle$

$$|\psi_n\rangle = N(\lambda) \left(|\phi_n\rangle + \sum_{k \neq n} C_{nk}(\lambda) |\phi_k\rangle \right)$$

$$\langle \psi_n | \psi_n \rangle = 1$$

power series expansion

- Perturbation theory evaluates the eigenvalues E_n , and the coefficients C_{nk} , as power series in λ

- coefficients

$$C_{nk}(\lambda) = \lambda C_{nk}^{(1)} + \lambda^2 C_{nk}^{(2)} + \dots$$

- eigenenergy

$$E_n = E_n^{(0)} + \lambda E_n^{(1)} + \lambda^2 E_n^{(2)} + \dots$$

Expansion

- The Schrodinger equation

$$\begin{aligned} & (H_0 + \lambda H_1) \left\{ |\phi_n\rangle + \sum_{k \neq n} \lambda C_{nk}^{(1)} |\phi_k\rangle + \sum_{k \neq n} \lambda^2 C_{nk}^{(2)} |\phi_k\rangle + \dots \right\} \\ &= \left(E_n^{(0)} + \lambda E_n^{(1)} + \lambda^2 E_n^{(2)} + \dots \right) \left\{ |\phi_n\rangle + \sum_{k \neq n} \lambda C_{nk}^{(1)} |\phi_k\rangle + \sum_{k \neq n} \lambda^2 C_{nk}^{(2)} |\phi_k\rangle + \dots \right\} \end{aligned}$$

- the 1st order

$$H_0 \sum_{k \neq n} \lambda C_{nk}^{(1)} |\phi_k\rangle + \lambda H_1 |\phi_n\rangle = E_n^{(0)} \sum_{k \neq n} \lambda C_{nk}^{(1)} |\phi_k\rangle + \lambda E_n^{(1)} |\phi_n\rangle$$

energy shift

$$H_0 \sum_{k \neq n} \lambda C_{nk}^{(1)} |\phi_k\rangle + \lambda H_1 |\phi_n\rangle = E_n^{(0)} \sum_{k \neq n} \lambda C_{nk}^{(1)} |\phi_k\rangle + \lambda E_n^{(1)} |\phi_n\rangle$$

$$E_k \sum_{k \neq n} C_{nk}^{(1)} |\phi_k\rangle + H_1 |\phi_n\rangle - E_n^{(0)} \sum_{k \neq n} C_{nk}^{(1)} |\phi_k\rangle = E_n^{(1)} |\phi_n\rangle$$

$$H_1 |\phi_n\rangle + \sum_{k \neq n} (E_k^{(0)} - E_n^{(0)}) C_{nk}^{(1)} |\phi_k\rangle = E_n^{(1)} |\phi_n\rangle$$

$$E_n^{(1)} \simeq \langle \phi_n | H_1 | \phi_n \rangle$$

- energy shift

$$\lambda E_n^{(1)} \simeq \langle \phi_n | \lambda H_1 | \phi_n \rangle$$

wavefunction change

- choose $m \neq n$

$$\langle \phi_m | H_1 | \phi_n \rangle + (E_m^{(0)} - E_n^{(0)}) C_{nm}^{(1)} = 0$$

$$C_{nm}^{(1)} = \frac{\langle \phi_m | H_1 | \phi_n \rangle}{E_n^{(0)} - E_m^{(0)}}$$

- wavefunction change

$$\lambda C_{nm}^{(1)} = \frac{\langle \phi_m | \lambda H_1 | \phi_n \rangle}{E_n^{(0)} - E_m^{(0)}}$$

$$|\psi_n\rangle = |\phi_n\rangle + \sum_{k \neq n} \frac{\langle \phi_k | \lambda H_1 | \phi_n \rangle}{E_n^{(0)} - E_k^{(0)}} |\phi_k\rangle$$

2nd order

- 2nd order

$$\begin{aligned} & H_0 \sum_{k \neq n} C_{nk}^{(2)} |\phi_k\rangle + H_1 \sum_{k \neq n} C_{nk}^{(1)} |\phi_k\rangle \\ &= E_n^{(2)} |\phi_n\rangle + E_n^{(1)} \sum_{k \neq n} C_{nk}^{(1)} |\phi_k\rangle + E_n^{(0)} \sum_{k \neq n} C_{nk}^{(2)} |\phi_k\rangle \end{aligned}$$

- energy change

$$H_0 \sum_{k \neq n} C_{nk}^{(2)} |\phi_k\rangle = \sum_{k \neq n} E_k^{(0)} C_{nk}^{(2)} |\phi_k\rangle$$

$$\begin{aligned} E_n^{(2)} &= \langle \phi_n | H_1 \sum_{k \neq n} C_{nk}^{(1)} |\phi_k\rangle = \sum_{k \neq n} C_{nk}^{(1)} \langle \phi_n | H_1 | \phi_k \rangle \\ &= \sum_{k \neq n} \frac{\langle \phi_n | H_1 | \phi_k \rangle \langle \phi_k | H_1 | \phi_n \rangle}{E_n^{(0)} - E_k^{(0)}} \\ &= \sum_{k \neq n} \frac{|\langle \phi_n | H_1 | \phi_k \rangle|^2}{E_n^{(0)} - E_k^{(0)}} \end{aligned}$$

Degenerate case

- When there are degenerate states, the perturbation breaks down $E_n = E_k$
- To solve this problem, we may diagonalize the perturbation in the degeneracy subspace

Example: 3-level system

- consider a model Hamiltonian

$$H_0 = \begin{pmatrix} \boxed{E^{(0)} & 0} & 0 \\ 0 & \boxed{E^{(0)}} & 0 \\ 0 & 0 & E_3^{(0)} \end{pmatrix} \quad \begin{matrix} \text{degeneracy} \\ \text{subspace} \end{matrix} \quad H_1 = \begin{pmatrix} h_{11} & h_{12} & h_{13} \\ h_{21} & h_{22} & h_{23} \\ h_{31} & h_{32} & h_{33} \end{pmatrix}$$

- Diagonalize H using a unitary transformation

$$UHU^\dagger = H_D$$

$$H_D = \begin{pmatrix} \boxed{E^{(0)} + \lambda w_1} & 0 & \lambda h_{13} \\ 0 & \boxed{E^{(0)} + \lambda w_2} & \lambda h_{23} \\ \lambda h_{31} & \lambda h_{32} & E_3^{(0)} + \lambda h_{33} \end{pmatrix}$$

$$w_{1,2} = \frac{h_{11} + h_{22}}{2} \pm \sqrt{\left(\frac{h_{11} - h_{22}}{2}\right)^2 + h_{12}h_{21}}$$

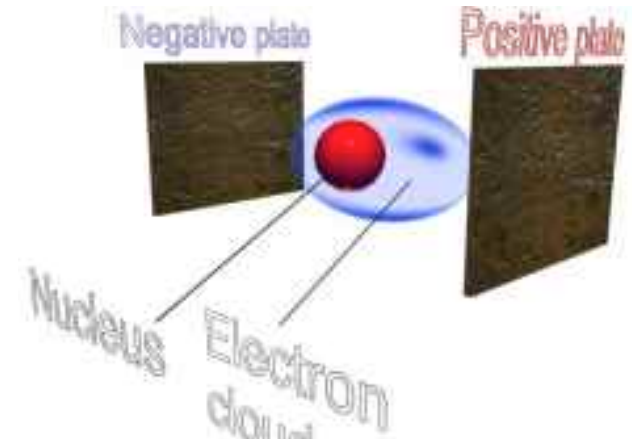
- then the degeneracy is lifted

The Stark effect

- Atom in the E-field

$$H_0 = \frac{p^2}{2\mu} - \frac{Ze^2}{4\pi\epsilon_0 r}$$

$$\lambda H_1 = e\mathcal{E}z$$



<http://www.sr.bham.ac.uk/xmm/fmc2.html>

- 1st order energy shift for ground state

$$E_{100}^{(1)} = e\mathcal{E} \langle \phi_{100} | z | \phi_{100} \rangle = e\mathcal{E} \int d^3r |\phi_{100}(r)|^2 z = 0$$

because of parity symmetry

2nd order shift

- 2nd order shift

$$E_{100}^{(2)} = e^2 \mathcal{E}^2 \left\{ \sum_{nlm} \frac{|\langle \phi_{nlm} | z | \phi_{100} \rangle|^2}{E_1^{(0)} - E_n^{(0)}} + \sum_k \frac{|\langle \phi_k | z | \phi_{100} \rangle|^2}{E_1^{(0)} - \frac{\hbar^2 k^2}{2m}} \right\}$$

$E > 0$ $|\phi_k\rangle$ unbound state
 $E < 0$ $|\phi_{nlm}\rangle$ bound state

$$E_{100}^{(2)} = e^2 \mathcal{E}^2 \sum_{E \neq E_1} \frac{|\langle \phi_E | z | \phi_{100} \rangle|^2}{E_1^{(0)} - E}$$

2nd order shift

- The upper-bound of energy shift

$$-E_{100}^{(2)} \leq e^2 \mathcal{E}^2 \frac{1}{E_2^{(0)} - E_1^{(0)}} \sum_{E \neq E_1} |\langle \phi_{100} | z | \phi_E \rangle|^2 = e^2 \mathcal{E}^2 \frac{1}{E_2^{(0)} - E_1^{(0)}} \sum_E |\langle \phi_{100} | z | \phi_E \rangle|^2$$

$$-E_{100}^{(2)} \leq e^2 \mathcal{E}^2 \frac{1}{E_2^{(0)} - E_1^{(0)}} \langle \phi_{100} | z^2 | \phi_{100} \rangle$$

$$\begin{aligned} \langle \phi_{100} | z^2 | \phi_{100} \rangle &= \langle \phi_{100} | x^2 | \phi_{100} \rangle = \langle \phi_{100} | y^2 | \phi_{100} \rangle \\ &= \frac{1}{3} \langle \phi_{100} | r^2 | \phi_{100} \rangle = a_0^2 \end{aligned}$$

$$E_2^{(0)} - E_1^{(0)} = -\frac{1}{2} m_e c^2 \alpha^2 \left(\frac{1}{4} - 1 \right) = \frac{3}{8} m_e c^2 \alpha^2$$

2nd order shift

- energy shift

$$-E_{100}^{(2)} \leq e^2 \mathcal{E}^2 \frac{8}{3} \frac{a_0^2}{m_e c^2 \alpha^2} = \frac{8}{3} (4\pi\epsilon_0 \mathcal{E}^2) a_0^3$$

$$-E_{100}^{(2)} = (\cos nt) (4\pi\epsilon_0 \mathcal{E}^2) a_0^3 = \frac{9}{4} (4\pi\epsilon_0 \mathcal{E}^2) a_0^3$$

n=2 state

- 4-level degeneracy

$$|\phi_{200}\rangle$$

$$|\phi_{211}\rangle$$

$$|\phi_{210}\rangle$$

$$|\phi_{2,1,-1}\rangle$$

- parity: $l=0$ even, $l=1$ odd

symmetry of H_1

- the perturbation commutes with L_z

$$[z, L_z] = 0$$

- matrix elements of H_1 involving states with different m are zero

$$\langle l, m | L_z H_1 - H_1 L_z | l', m' \rangle = 0 \quad (m - m') \langle l, m | H_1 | l', m' \rangle = 0$$

- Only the matrix elements need to be considered

$$\langle \phi_{200} | H_1 | \phi_{210} \rangle \quad \langle \phi_{200} | H_1 | \phi_{200} \rangle \quad \langle \phi_{210} | H_1 | \phi_{210} \rangle$$

degenerate case

- With the technique of degenerate perturbation

$$e\mathcal{E} \begin{pmatrix} \langle \phi_{200} | z | \phi_{200} \rangle & \langle \phi_{200} | z | \phi_{210} \rangle \\ \langle \phi_{210} | z | \phi_{200} \rangle & \langle \phi_{210} | z | \phi_{210} \rangle \end{pmatrix} \begin{pmatrix} \alpha_1 \\ \alpha_2 \end{pmatrix} = E^{(1)} \begin{pmatrix} \alpha_1 \\ \alpha_2 \end{pmatrix}$$

$$e\langle \phi_{200} | z | \phi_{210} \rangle = \left[\int \frac{e^{-r/a_0}}{(2a_0)^3} \frac{2r}{\sqrt{3}a_0} \left(1 - \frac{r}{2a_0}\right) rr^2 dr \right] \left[\int d\Omega Y_{00}^* \sqrt{\frac{4\pi}{3}} Y_{10} Y_{10} \right] = -3a_0$$

$$\langle \phi_{200} | H_1 | \phi_{200} \rangle = \langle \phi_{210} | H_1 | \phi_{210} \rangle = 0$$

energy shift for $m=0$ states

- The eigenvalues
$$-3e\mathcal{E}a_0 \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix} \begin{pmatrix} \alpha_1 \\ \alpha_2 \end{pmatrix} = E^{(1)} \begin{pmatrix} \alpha_1 \\ \alpha_2 \end{pmatrix}$$

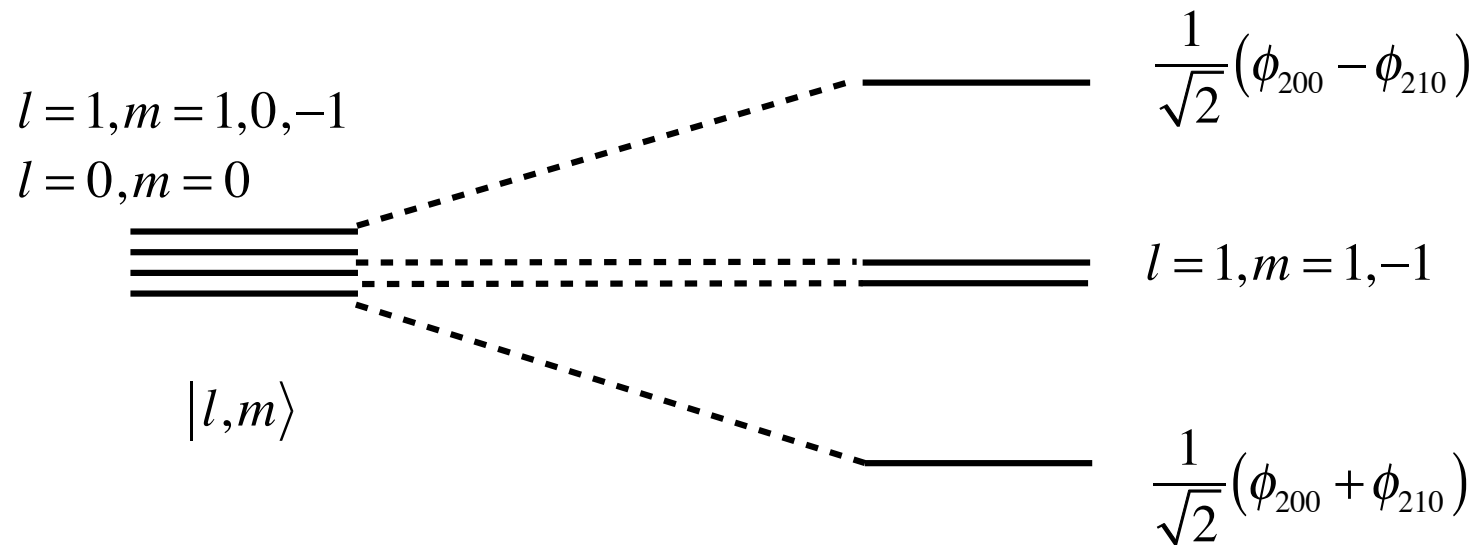
$$E^{(1)} = \pm 3e\mathcal{E}a_0$$

- The eigenstates

$$\frac{1}{\sqrt{2}} \begin{pmatrix} 1 \\ -1 \end{pmatrix}$$

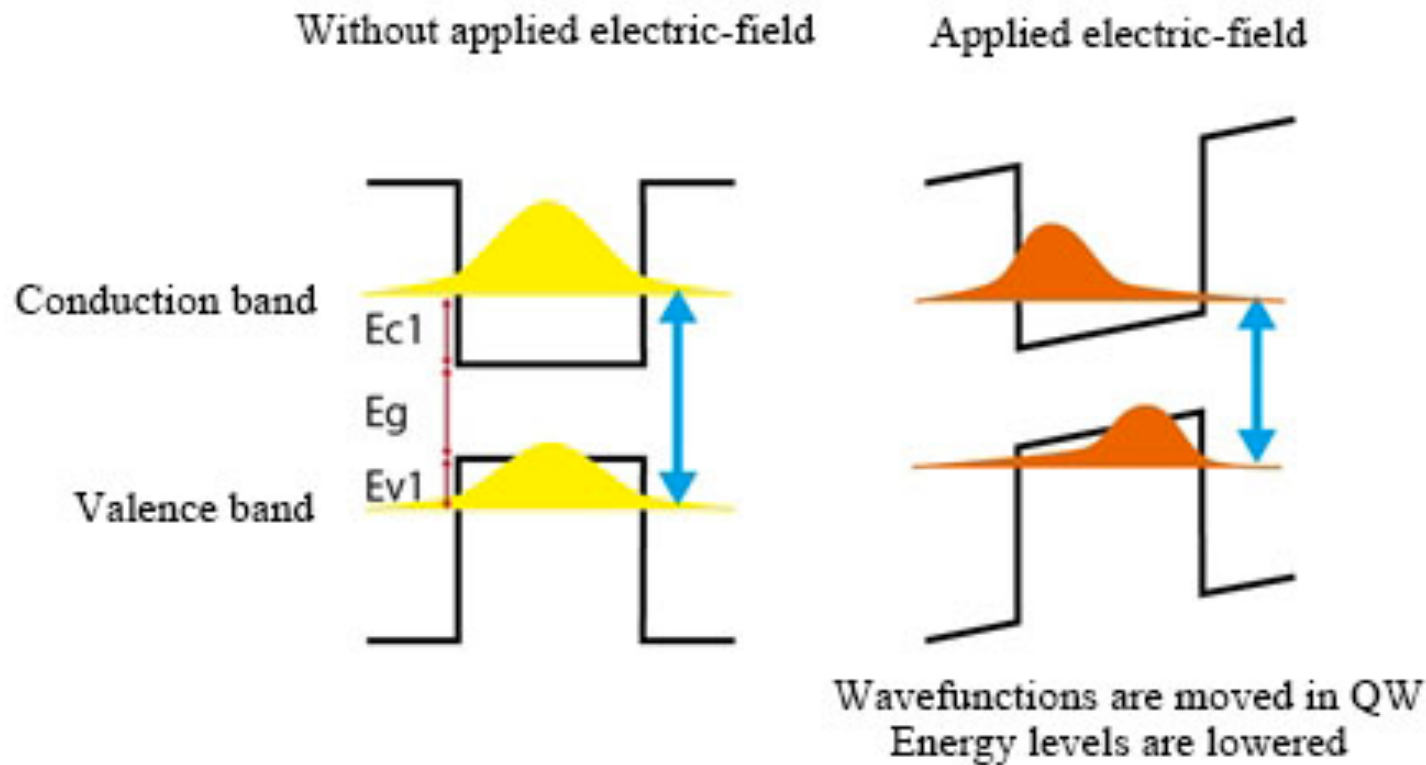
$$\frac{1}{\sqrt{2}} \begin{pmatrix} 1 \\ 1 \end{pmatrix}$$

Stark shift for n=2 states



- The electron may tunnel out if E is strong

Stark effect in Q well



<http://pweb.cc.sophia.ac.jp/shimolab/html-e/qcse-e.html>

Fine structure

- The fine structure of hydrogen atom contains the relativistic correction and SOI
- The relativistic correction for the kinetic energy is

$$K = \sqrt{p^2 c^2 + m_e^2 c^4} - m_e c^2 = \frac{p^2}{2m_e} - \frac{1}{8} \frac{p^4}{m_e^3 c^2} + \dots$$

$$H_1 = -\frac{1}{8} \frac{p^4}{m_e^3 c^2}$$

- spin orbital interaction(SOI)

$$H_2 = \frac{Ze^2}{4\pi\epsilon_0} \frac{1}{2m_e^2 c^2 r^3} \mathbf{S} \cdot \mathbf{L}$$

Relativistic correction

- 1st order energy correction $H_0 = \frac{p^2}{2m_e} - \frac{Ze^2}{4\pi\epsilon_0 r}$

$$\begin{aligned}\langle \phi_{nlm} | H_1 | \phi_{nlm} \rangle &= -\frac{1}{2m_e c^2} \langle \phi_{nlm} | \left(\frac{p^2}{2m_e} \right)^2 | \phi_{nlm} \rangle \\ &= -\frac{1}{2m_e c^2} \langle \phi_{nlm} | \left(H_0 + \frac{Ze^2}{4\pi\epsilon_0 r} \right)^2 | \phi_{nlm} \rangle \\ &= -\frac{1}{2m_e c^2} \langle \phi_{nlm} | \left(E_n + \frac{Ze^2}{4\pi\epsilon_0 r} \right)^2 | \phi_{nlm} \rangle \\ &= -\frac{1}{2m_e c^2} \left[E_n^2 + 2E_n \frac{Ze^2}{4\pi\epsilon_0} \langle \phi_{nlm} | \frac{1}{r} | \phi_{nlm} \rangle + \left(\frac{Ze^2}{4\pi\epsilon_0} \right)^2 \langle \phi_{nlm} | \frac{1}{r^2} | \phi_{nlm} \rangle \right]\end{aligned}$$

1st-order energy shift

$$\begin{aligned}
 \langle \phi_{nlm} | H_1 | \phi_{nlm} \rangle &= -\frac{1}{2m_e c^2} \left[E_n^2 + 2E_n \frac{Ze^2}{4\pi\epsilon_0} \langle \phi_{nlm} | \frac{1}{r} | \phi_{nlm} \rangle + \left(\frac{Ze^2}{4\pi\epsilon_0} \right)^2 \langle \phi_{nlm} | \frac{1}{r^2} | \phi_{nlm} \rangle \right] \\
 &= -\frac{1}{2m_e c^2} \left[\left(\frac{m_e c^2 \alpha^2 Z^2}{2 n^2} \right)^2 + 2 \frac{m_e c^2 \alpha^2 Z^2}{2 n^2} \left(\frac{Ze^2}{4\pi\epsilon_0} \right) \frac{Z}{a_0 n^2} + \left(\frac{Ze^2}{4\pi\epsilon_0} \right)^2 \frac{2Z^2}{a_0^2 n^3 (2l+1)} \right] \\
 &= -\frac{1}{2m_e c^2} \left[\left(\frac{m_e c^2 \alpha^2 Z^2}{2 n^2} \right)^2 - 2 \frac{m_e c^2 \alpha^2 Z^4}{2 n^4} (\alpha \hbar c) \left(\frac{\alpha m_e c}{\hbar} \right) + (\alpha \hbar c)^2 \left(\frac{\alpha m_e c}{\hbar} \right)^2 \frac{2Z^4}{n^3 (2l+1)} \right] \\
 &= -\frac{m_e c^2 \alpha^4 Z^4}{2 n^4} \left[\frac{1}{4} - 1 + \frac{2n}{2l+1} \right] = -\frac{1}{2} m_e c^2 \frac{\alpha^4 Z^4}{n^4} \left[\frac{2n}{2l+1} - \frac{3}{4} \right]
 \end{aligned}$$

$$\langle \phi_{nlm} | \frac{1}{r} | \phi_{nlm} \rangle = \frac{Z}{a_0 n^2} \qquad \frac{e^2}{4\pi\epsilon_0} = \alpha \hbar c$$

$$\langle \phi_{nlm} | \frac{1}{r^2} | \phi_{nlm} \rangle = \frac{2Z^2}{a_0^2 n^3 (2l+1)} \qquad a_0 = \frac{\alpha m_e c}{\hbar}$$

$$m_e c^2 \sim 0.5 \text{ MeV} \qquad \alpha^4 \sim \left(\frac{1}{137} \right)^4 \sim 3 \times 10^{-9} \qquad \Delta E_1 \sim 10^{-4} \text{ eV}$$

Spin-orbital coupling

- To calculate the SOI, we can use the result for addition of angular momenta

$$\mathbf{J} = \mathbf{L} + \mathbf{S}$$

$$J_z = L_z + S_z$$

$$\mathbf{S} \cdot \mathbf{L} = \frac{1}{2}(J^2 - L^2 - S^2)$$

- The possible eigenstates

$$|l, m\rangle \otimes \left| \frac{1}{2}, \frac{1}{2} \right\rangle$$

→

$$\left| l + \frac{1}{2}, m + \frac{1}{2} \right\rangle$$

$$j = l + \frac{1}{2}$$

parallel

$$\left| l - \frac{1}{2}, m + \frac{1}{2} \right\rangle$$

anti-parallel

$$j = l - \frac{1}{2}$$

Eigenvalues

- The eigenstates gives different SOI

- Parallel $\left|l + \frac{1}{2}, m + \frac{1}{2}\right\rangle$

$$\begin{aligned}\mathbf{S} \cdot \mathbf{L} \left|l + \frac{1}{2}, m + \frac{1}{2}\right\rangle &= \frac{1}{2} (J^2 - L^2 - S^2) \left|l + \frac{1}{2}, m + \frac{1}{2}\right\rangle \\ &= \frac{1}{2} \hbar^2 \left[\left(l + \frac{1}{2}\right) \left(l + \frac{3}{2}\right) - l(l+1) - \frac{1}{2} \frac{3}{2} \right] \left|l + \frac{1}{2}, m + \frac{1}{2}\right\rangle \\ &= \frac{1}{2} \hbar^2 l \left|l + \frac{1}{2}, m + \frac{1}{2}\right\rangle\end{aligned}$$

- Anti-parallel $\left|l - \frac{1}{2}, m + \frac{1}{2}\right\rangle$

$$\begin{aligned}\mathbf{S} \cdot \mathbf{L} \left|l - \frac{1}{2}, m + \frac{1}{2}\right\rangle &= \frac{1}{2} (J^2 - L^2 - S^2) \left|l - \frac{1}{2}, m + \frac{1}{2}\right\rangle \\ &= \frac{1}{2} \hbar^2 \left[\left(l - \frac{1}{2}\right) \left(l + \frac{1}{2}\right) - l(l+1) - \frac{1}{2} \frac{3}{2} \right] \left|l - \frac{1}{2}, m + \frac{1}{2}\right\rangle \\ &= -\frac{1}{2} \hbar^2 (l+1) \left|l - \frac{1}{2}, m + \frac{1}{2}\right\rangle\end{aligned}$$

SOI energy shift

- The energy shift for the 1st order

$$\begin{aligned}
 \langle \phi_{jm,l} | H_2 | \phi_{jm,l} \rangle &= \frac{Ze^2}{4\pi\epsilon_0} \frac{1}{2m_e^2 c^2} \langle \phi_{jm,l} | (\mathbf{S} \cdot \mathbf{L}) \frac{1}{r^3} | \phi_{jm,l} \rangle \\
 &= \frac{Ze^2}{4\pi\epsilon_0} \frac{1}{2m_e^2 c^2} \frac{\hbar^2}{2} \left\{ \begin{array}{c} l \\ -l-1 \end{array} \right\} \langle \phi_{jm,l} | \frac{1}{r^3} | \phi_{jm,l} \rangle \\
 &= \frac{Ze^2}{4\pi\epsilon_0} \frac{1}{2m_e^2 c^2} \frac{\hbar^2}{2} \left\{ \begin{array}{c} l \\ -l-1 \end{array} \right\} \frac{Z^3}{a_0^3} \frac{2}{n^3 l(l+1)(2l+1)} \\
 &= (\alpha \hbar c) \left(\frac{m_e c \alpha}{\hbar} \right)^3 \frac{\hbar^2 Z^4}{2m_e^2 c^2} \frac{\left\{ \begin{array}{c} l \\ -l-1 \end{array} \right\}}{n^3 l(l+1)(2l+1)} \\
 &= \frac{1}{2} m_e c^2 Z^4 \alpha^4 \frac{\left\{ \begin{array}{c} l \\ -l-1 \end{array} \right\}}{n^3 l(l+1)(2l+1)} = \frac{1}{2} m_e c^2 Z^4 \alpha^4 \frac{1}{n^3 (2l+1)} \left\{ \begin{array}{c} 1/(l+1) \\ -1/l \end{array} \right\} \\
 &= \pm \frac{1}{2} m_e c^2 Z^4 \alpha^4 \frac{1}{n^3 (2l+1) \left(j + \frac{1}{2} \right)} \quad \Delta E_2 \sim 10^{-4} \text{ eV}
 \end{aligned}$$

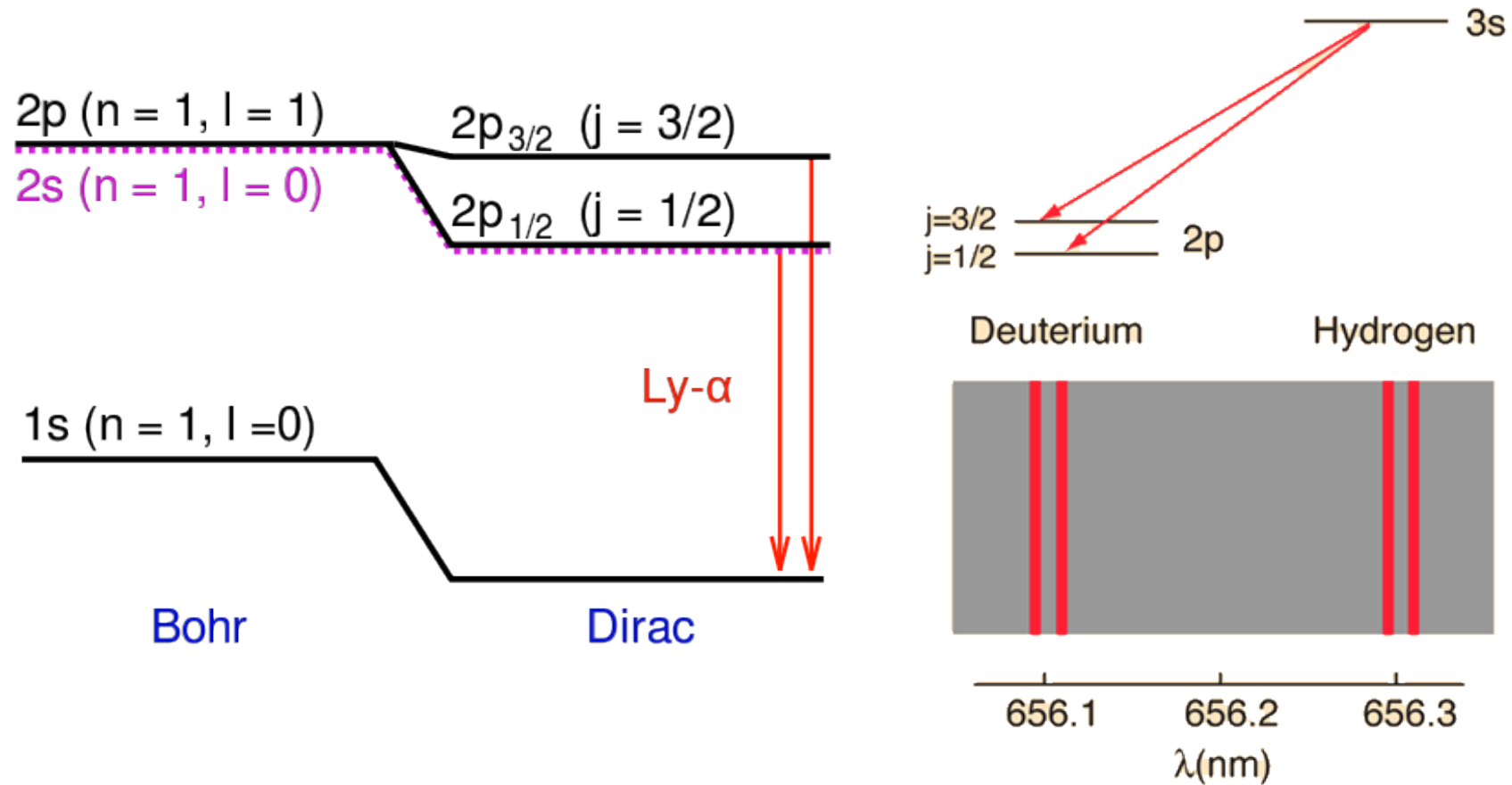
$j = l + \frac{1}{2}$
 $j = l - \frac{1}{2}$

Fine structure

- Total energy shift combining relativistic correction and SOI

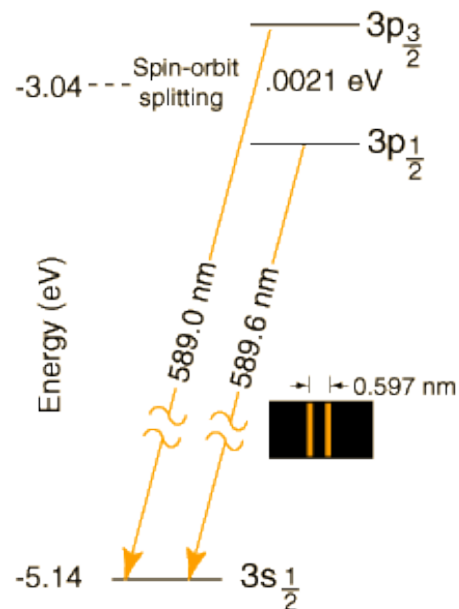
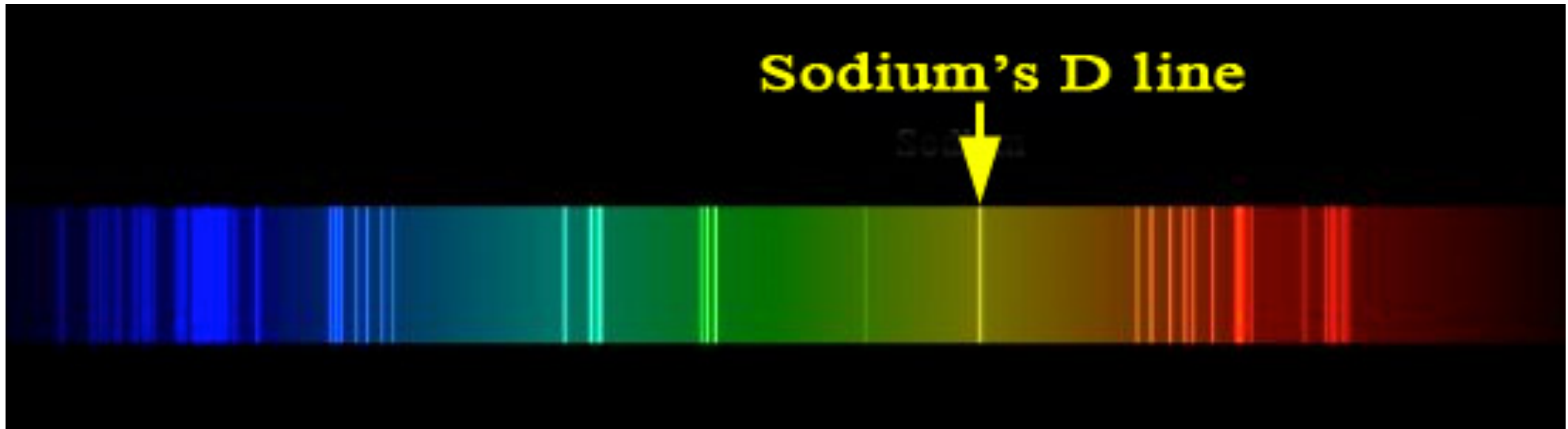
$$\begin{aligned}
 \langle \phi_{jm,l} | H_1 + H_2 | \phi_{jm,l} \rangle &= -\frac{m_e c^2 \alpha^4 Z^4}{2 n^4} \left[\frac{2n}{2l+1} - \frac{3}{4} \right] \pm \frac{1}{2} m_e c^2 Z^4 \alpha^4 \frac{1}{n^3 (2l+1)(j+1/2)} \\
 &= -\frac{1}{2} m_e c^2 Z^4 \alpha^4 \frac{1}{n^3} \left[\frac{1}{(2l+1)} \left(\mp \frac{1}{j+1/2} + 2 \right) - \frac{3}{4n} \right] \\
 &= -\frac{1}{2} m_e c^2 Z^4 \alpha^4 \frac{1}{n^3} \left[\frac{1}{(2l+1)(j+1/2)} \binom{2j}{2j+2} - \frac{3}{4n} \right] \\
 &= -\frac{1}{2} m_e c^2 Z^4 \alpha^4 \frac{1}{n^3} \left[\frac{1}{j+1/2} - \frac{3}{4n} \right]
 \end{aligned}$$

Fine structure



<http://hyperphysics.phy-astr.gsu.edu/hbase/quantum/hydfin.html>

Sodium doublet



<http://hyperphysics.phy-astr.gsu.edu/hbase/quantum/hydfin.html>

Hyperfine structure

- The magnetic moment of nucleus produces a very tiny magnetic field to the electron
- The nuclear magnetic moment is related to its spin

$$\mathbf{M} = \frac{Ze g_N}{2M_N} \mathbf{I}$$

- the dipole field is
$$\mathbf{B} = \frac{\mu_0}{4\pi} \left(\frac{3\mathbf{r}(\mathbf{r} \cdot \mathbf{M}) - r^2 \mathbf{M}}{r^5} + \frac{8\pi}{3} \mathbf{M} \delta(\mathbf{r}) \right)$$

- The energy is

$$H_{hf} = -\mathbf{m} \cdot \mathbf{B} = \frac{e}{2m_e} (\mathbf{L} + 2\mathbf{S}) \cdot \mathbf{B}$$

$$\mathbf{m} = \frac{e}{2m_e} \mathbf{L} + \frac{eg}{2m_e} \mathbf{S}$$

perturbation

- The hyperfine perturbation

$$\begin{aligned} H_{hf} &= \frac{e}{2m_e} \frac{Zeg_N}{2M_N} \frac{\mu_0}{4\pi} \left(\frac{3\mathbf{r}(\mathbf{r} \cdot \mathbf{I}) - r^2\mathbf{I}}{r^5} + \frac{8\pi}{3}\mathbf{I}\delta(\mathbf{r}) \right) \cdot (\mathbf{L} + 2\mathbf{S}) \\ &= \frac{Ze^2}{4\pi\epsilon_0} \frac{g_N}{4M_N m_e c^2} \left(\frac{3\mathbf{r}(\mathbf{r} \cdot \mathbf{I}) - r^2\mathbf{I}}{r^5} + \frac{8\pi}{3}\mathbf{I}\delta(\mathbf{r}) \right) \cdot (\mathbf{L} + 2\mathbf{S}) \end{aligned}$$

- If $\mathbf{L} = 0$

$$H_{hf} = \frac{Ze^2}{4\pi\epsilon_0} \frac{g_N}{2M_N m_e c^2} \left(\frac{3(\mathbf{r} \cdot \mathbf{I})(\mathbf{r} \cdot \mathbf{S}) - r^2\mathbf{I}}{r^5} + \frac{8\pi}{3}(\mathbf{I} \cdot \mathbf{S})\delta(\mathbf{r}) \right).$$

- The energy shift

$$\langle \phi_{n00} | H_{hf} | \phi_{n00} \rangle = \frac{Ze^2}{4\pi\epsilon_0} \frac{g_N}{2M_N m_e c^2} \frac{8\pi}{3} \langle \phi_{n00} | (\mathbf{I} \cdot \mathbf{S}) \delta(\mathbf{r}) | \phi_{n00} \rangle$$

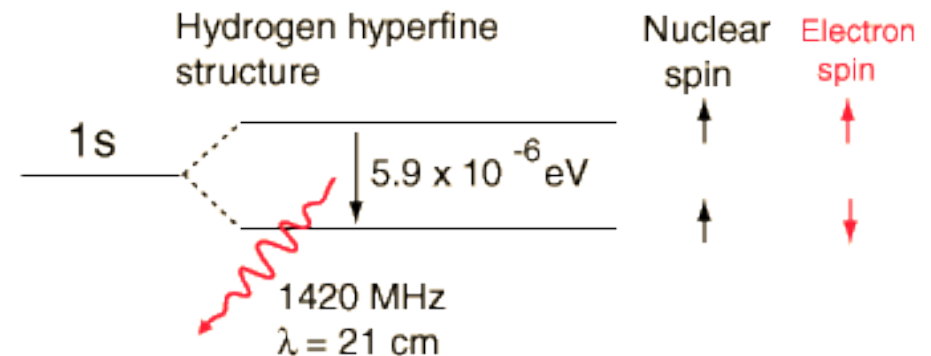
energy shift

- For H-atom, $n=1$

$$\begin{aligned} \langle \phi_{n00} | H_{hf} | \phi_{n00} \rangle &= \frac{8\pi}{3} \frac{g_N}{2M_N m_e c^2} (\alpha \hbar c) \left(\frac{m_e \alpha c}{\hbar} \right)^3 \frac{1}{\pi} (\mathbf{I} \cdot \mathbf{S}) \\ &= \frac{4}{3} m_e c^2 \alpha^4 g_N \frac{m_e}{M_N} \frac{\langle \mathbf{I} \cdot \mathbf{S} \rangle}{\hbar^2} \end{aligned}$$

- hyperfine splitting $\langle \mathbf{I} \cdot \mathbf{S} \rangle = \frac{1}{2} \langle F^2 - I^2 - S^2 \rangle = \left\{ \begin{array}{l} 1/4 \\ -3/4 \end{array} \right\} \quad \mathbf{F} = \mathbf{I} + \mathbf{S}$

$$\begin{aligned} \Delta E_{hf} &= \frac{4}{3} m_e c^2 \alpha^4 g_N \frac{m_e}{M_N} \left(\frac{1}{4} - \left(\frac{3}{4} \right) \right) \\ &= \frac{4}{3} m_e c^2 \alpha^4 g_N \frac{m_e}{M_N} \end{aligned}$$



energy splitting

- Put in values $g_N = g_p = 5.56$ $\frac{m_e}{M_p} = \frac{1}{1830}$

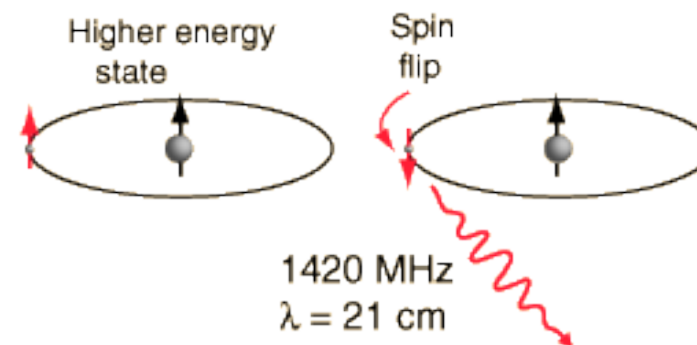
- energy splitting

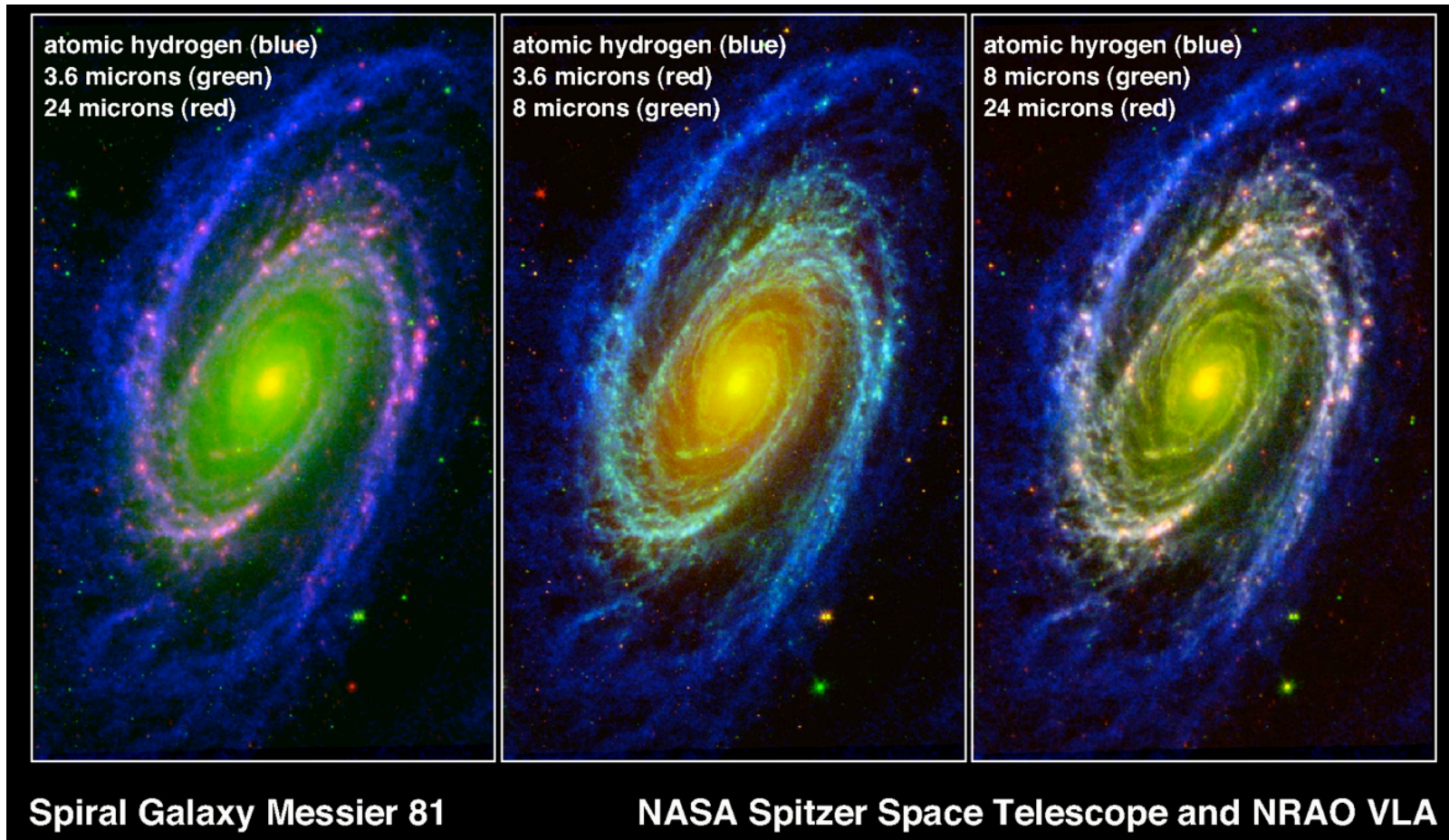
$$\begin{aligned}\Delta E_{hf} &= \frac{4}{3}(0.5\text{M eV})\left(\frac{1}{137}\right)^4 (5.56)\left(\frac{1}{1830}\right) \\ &= 5.7 \times 10^{-6} \text{ eV} \\ &\sim (1420 \text{ MHz})h\end{aligned}$$

- transition between the splitting

$$\lambda = \sim 21 \text{ cm}$$

used in radio astronomy





Gallery

image of M81 made with the VLA for the THINGS

Very Large Array(VLA) National Radio Astronomy Observatory (NRAO)



